University of California at San Diego – Department of Physics – Prof. John McGreevy

## Physics 215A QFT Fall 2022 Assignment 5

Due 11:00am Thursday, October 27, 2022

1. Brain-warmer: the identity does nothing twice. Check our relativistic state normalization by squaring the expression for the identity in the 1-particle sector:

$$\mathbb{1}_1^2 \stackrel{!}{=} \mathbb{1}_1 = \int \frac{\mathrm{d}^d p}{2\omega_{\vec{p}}} \left| \vec{p} \right\rangle \left\langle \vec{p} \right|.$$

## 2. Even more about QFT in 0 + 0 dimensions.

In this problem we return to the simplest scalar QFT, namely a single integral with quartic action.

(a) By a change of integration variable show that

$$Z = \int_{-\infty}^{\infty} dq \ e^{-S(q)}$$

with  $S(q) = \frac{1}{2}m^2q^2 + gq^4$  is of the form

$$Z = rac{1}{g^{1/4}} \mathcal{Z} \left( m^2 / \sqrt{g} 
ight).$$

This means you can make your life easier by setting g = 1, without loss of generality.

- (b) Convince yourself (*e.g.* with Mathematica) that the integral really is expressible as a Bessel function.
- (c) It would be nice to find a better understanding for why the partition function of (0+0)-dimensional  $\phi^4$  theory is a Bessel function. Find a Schwinger-Dyson equation for this system which has the form of Bessel's equation for

$$K(x^2/a) \equiv e^{-x^2/a} (x^2)^{-1/4} \mathcal{Z}(x)$$

for some constant *a*. (If you get stuck I can tell you what function to choose for the 'anything' in the S-D equation.)

(d) Make a plot of the perturbative approximations to the 'Green function'  $G \equiv \langle q^2 \rangle$  as a function of g, truncated at orders 1 through 6 or so. Plot them against the exact answer.

(e) (Bonus problem) Show that  $c_{n+1} \sim -\frac{2}{3}nc_n$  at large n, where  $c_n$  is the coefficient of  $g^n$  in either G or Z (by brute force or by cleverness).

## 3. The vacuum is a fluid with $p = -\rho$ . [Bonus problem]

We said in lecture that the vacuum energy density  $\rho$  gravitates and that, when positive, its effect is to cause space to inflate – to expand exponentially in time. An important aspect of this phenomenon is that the vacuum fluctuations produce not only an energy density, but a *pressure*,  $p = T_i^i$  (no sum on *i*), of the form  $p = -\rho$ , which is negative for  $\rho > 0$ . The vacuum therefore acts as a perfect fluid with  $P = -\rho$ . (The stress tensor for a perfect fluid in terms of its velocity field  $u^{\mu}$  takes the form  $T^{\mu\nu} = (p + \rho)u^{\mu}u^{\nu} + pg^{\mu\nu}$ , so in a frame with  $u^{\mu} = (1, \vec{0}^{\mu})$ ,  $T_0^0 = \rho, T_i^i = P$ .) Solving Einstein's equations with such a source produces an inflating universe. In this problem we show that this is is the case.

(a) Show that the energy-momentum tensor for a free relativistic scalar field  $(S[\phi] = \int d^D x \sqrt{g} \mathcal{L}, \mathcal{L} = \frac{1}{2} g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi - \frac{m^2}{2} \phi^2)$  takes the form

$$T_{\mu\nu} = a\partial_{\mu}\phi\partial_{\nu}\phi - bg_{\mu\nu}\mathcal{L}$$

with some constants a, b.

(b) Reproduce the formal expression for the vacuum energy

$$\langle 0|\mathbf{H}|0\rangle = V \int \mathrm{d}^d k \frac{1}{2} \hbar \omega_{\vec{k}}$$

using the two point function

$$\langle 0 | \phi(x)^2 | 0 \rangle = \langle 0 | \phi(0) \phi(0) | 0 \rangle = \lim_{\vec{x}, t \to 0} \langle 0 | \phi(x) \phi(0) | 0 \rangle$$

and its derivatives. (V is the volume of space.)

(We will learn to draw this amplitude as a Feynman diagram which is a circle (a line connecting a point to itself).)

(c) Show that the vacuum expectation value of the pressure

 $\langle 0|T_{ii}|0\rangle$ 

(no sum on i) gives the same answer, up to a sign.

[Hints: You'll find a quite different looking integral from the vacuum energy. Use rotation invariance of the vacuum to simplify the answer. The claim is that however you regulate the integral vacuum pressure and  $\frac{1}{2} \int d^d k \omega_k$ , you'll get the same answer. A convenient regulator is *dimensional regularization*: treat the dimension d as an arbitrary complex number.]

(d) Show that the resulting vacuum energy momentum tensor  $(T_{00} = \rho, T_{ii} = -\rho)$  (no sum on *i*)) is the same as the contribution to the energy-momentum tensor from an action of the form

$$S_{\rm cc} = \int d^D x \sqrt{g} \Lambda$$

where  $\Lambda$  is a constant (the cosmological constant).

- (e) Argue that  $p = -\rho$  is required in order that the vacuum energy does not specify a preferred rest frame.
- 4. Non-Abelian currents. On a previous homework, we studied a complex scalar field. Now, we make a big leap to *two* complex scalar fields,  $\Phi_{\alpha=1,2}$ , with

$$S[\Phi_{\alpha}] = \int d^{d}x dt \left(\frac{1}{2}\partial_{\mu}\Phi_{\alpha}^{\star}\partial^{\mu}\Phi_{\alpha} - V\left(\Phi_{\alpha}^{\star}\Phi_{\alpha}\right)\right)$$

Consider the objects

$$Q^{i} \equiv \frac{1}{2} \int d^{d}x \mathbf{i} \left( \Pi^{\dagger}_{\alpha} \sigma^{i}_{\alpha\beta} \Phi^{\dagger}_{\beta} \right) + h.c.$$

where  $\sigma^{i=1,2,3}$  are the three Pauli matrices.

- (a) What symmetries do these charges generate (*i.e.* how do the fields transform)? Show that they are symmetries of S.
- (b) If you want to, show that  $[Q^i, H] = 0$ , where H is the Hamiltonian.
- (c) Evaluate  $[Q^i, Q^j]$ . Hence, non-Abelian.
- (d) To complete the circle, find the Noether currents  $J^i_{\mu}$  associated to the symmetry transformations you found in part 4a. Show that the resulting charge  $Q^i = \int d^d x J^i_0$  agrees with our starting point.
- (e) Generalize to the case of N scalar fields.
- 5. Combinatorics from 0-dimensional QFT. [This is a more sophisticated bonus problem. I will not post the solutions of this problem until later. If you have a hard time with it now, please try again in a week.]

Catalan numbers  $C_n = \frac{(2n)!}{n!(n+1)!}$  arise as the answer to many combinatorics problems (beware: there is some disagreement in the literature about whether this is  $C_n$  or  $C_{n+1}$ ).

One such problem is: count random walks on a 1d chain with 2n steps that start at 0 and end at 0 without crossing 0 in between.



Another such problem is: in how many ways can 2n (distinguishable) points on a circle be connected by chords that do not intersect within the circle.



Consider a zero-dimensional QFT with the following Feynman rules:

- There are two fields h and l.
- There is an  $\sqrt{t}h^2l$  vertex in terms of a coupling t.
- The bare l propagator is 1.
- The bare h propagator is 1.
- All diagrams can be drawn on a piece of paper without crossing.
- An annoying extra rule: All the l propagators must be on one side of the h propagators<sup>1</sup>.
- There are no loops of h.

The last two rules can be realized from a lagrangian by introducing a large N (below).

(a) Show that the full two-point green's function for h is

$$G(t) = \sum_{n} t^{n} C_{n}$$

the generating function of Catalan numbers.

- (b) Let  $\Sigma(t)$  be the sum of diagrams with two *h* lines sticking out that may not be divided into two parts by cutting a single intermediate line. (This property is called 1PI (one-particle irreducible), and  $\Sigma$  is called the "1PI self-energy of *h*". We'll use this manipulation all the time later on.) Show that  $G(t) = \frac{1}{1-\Sigma(t)}$ .
- (c) Argue by diagrams for the equation (sometimes this is also called a Schwinger-Dyson equation)



<sup>&</sup>lt;sup>1</sup>You'll see in part 5f how to justify this.

where  $\Sigma$  is the 1PI self-energy of h.

- (d) Solve this equation for the generating function G(t).
- (e) If you are feeling ambitious, add another coupling  $N^{-1}$  which counts the crossings of the *l* propagators. The resulting numbers can be called Touchard-Riordan numbers.
- (f) How to realize the no-crossings rule? Consider

$$L = \frac{\sqrt{t}}{\sqrt{N}} l_{\alpha\beta} h_{\alpha} h_{\beta} + \sum_{\alpha,\beta} l_{\alpha\beta}^2 + \sum_{\alpha} h_{\alpha}^2$$

where  $\alpha, \beta = 1 \cdots N$ . By counting index loops, show that the dominant diagrams at large N are the ones we kept above. Hint: to keep track of the factors of N, introduce ('t Hooft's) double-line notation: since l is a matrix, its propagator looks like:  $\beta - - - - - - \alpha$ , while the h propagator is just one index line  $\alpha_{---\alpha}$ , and the vertex is \_\_\_\_!\_\_\_. If you don't like my ascii diagrams, here are the respective pictures:  $\langle l_{\alpha\beta}l_{\alpha\beta}\rangle = \langle A_{\alpha\beta}h_{\alpha}\rangle = \langle A_{\alpha\beta}h_{\alpha}\rangle = \langle A_{\alpha\beta}h_{\alpha\beta}\rangle = \langle A_{\alpha\beta}h_{\alpha\beta}\rangle$ 

closed index loop gives a factor of N.

- (g) Use properties of Catalan numbers to estimate the size of non-perturbative effects in this field theory.
- (h) There are many other examples like this. Another similar one is the relationship between symmetric functions and homogeneous products. A more different one is the enumeration of planar graphs. For that, see BIPZ.