

Physics 215A QFT Fall 2023 Assignment 7

Due 11:00am Monday, November 20, 2023

1. Brain-warmer: Wick example.

For a real scalar field, verify by hand Wick's prediction for the difference

$$\mathcal{T}(\phi(x_1)\phi(x_2)\phi(x_3)) - : \phi(x_1)\phi(x_2)\phi(x_3) :$$

2. Brain-warmers: Feynman rules.

Consider the field theory with action

$$S[\phi] = \int d^{d+1}x \left(\frac{1}{2}(\partial_\mu\phi\partial^\mu\phi - m^2\phi^2) - \frac{g}{3!}\phi^3 \right).$$

- Briefly state the Feynman rules in position space, emphasizing the differences from the ϕ^4 theory.
- Draw the diagrams that correct the position-space two-point function at order g^2 .
- Find the symmetry factor for these diagrams and verify them directly.

3. Particle creation by an external source.

Compare this problem with problem 6 on HW06.

Consider the Hamiltonian

$$H = H_0 + \int d^3x (-j(t, \vec{x})\phi(x))$$

where H_0 is the free Klein-Gordon Hamiltonian, ϕ is the Klein-Gordon field, and j is a c-number scalar function.

- Show that the probability that the source creates *no* particles is given by

$$P(0) = |\langle 0 | \mathcal{T}\{e^{+i\int d^4x j(x)\phi_I(x)}\} | 0 \rangle|^2.$$

- Evaluate the term in $P(0)$ of order j^2 , and show that $P(0) = 1 - \lambda + \mathcal{O}(j^4)$ where

$$\lambda = \int \frac{d^3p}{2E_p} |\tilde{j}(p)|^2.$$

We will show below that $\lambda = \langle N \rangle$ is the mean number of particles created by the source.

- (c) Represent the term computed in part 3b as a Feynman diagram. Now represent the whole perturbation series for $P(0)$ in terms of Feynman diagrams. (Hint: you have done this calculation already.) Show that this series exponentiates, so that it can be summed exactly $P(0) = e^{-\lambda}$.
- (d) On the next problem set, after learning about the notion of final-state phase space, we'll find the probability for the source to create any number of particles.

4. **Propagator corrections in a solvable field theory.**

Consider a theory of a scalar field in D dimensions with action

$$S = S_0 + S_1$$

where

$$S_0 = \int d^D x \frac{1}{2} (\partial_\mu \phi \partial^\mu \phi - m_0^2 \phi^2)$$

and

$$S_1 = - \int d^D x \frac{1}{2} \delta m^2 \phi^2 .$$

We have artificially decomposed the mass term into two parts. We will do perturbation theory in small δm^2 , treating S_1 as an 'interaction' term. We wish to show that the organization of perturbation theory that we've seen lecture will correctly reassemble the mass term.

- (a) Write down all the Feynman rules for this perturbation theory.
- (b) Determine the 1PI two-point function in this model, defined by

$$-i\Sigma \equiv \sum (\text{all 1PI diagrams with two nubbins}).$$

- (c) Show that the (geometric) summation of the propagator corrections correctly produces the propagator that you would have used had we not split up $m_0^2 + \delta m^2$.

5. **Wick's theorem from Schwinger-Dyson equations.** [Bonus problem] Study the derivation of Wick's theorem from the Schwinger-Dyson equation for the n -point function of a free scalar field on page 81 of Schwartz' book.

6. **A background field.** [This is a bonus problem.]

Consider the following action for a real scalar field Φ :

$$S[\Phi] = \int d^{d+1}x \frac{1}{2} (\partial_\mu \Phi \partial^\mu \Phi - m^2 \Phi^2 - g\phi(x)\Phi^2) .$$

The last term here is a cubic coupling between ϕ and Φ . But here we will treat $\phi(x)$ as a fixed background field (analogous to $j(x)$ on previous problems) which acts as a spacetime-dependent mass for the dynamical field Φ .

- (a) Show that the two-point Green's function, $G(x, y) \equiv \langle \Omega | \mathcal{T} \Phi(x) \Phi(y) | \Omega \rangle$, satisfies the Schwinger-Dyson equation

$$-\mathbf{i}\delta^{d+1}(x-y) = (\partial^2 + m^2 + g\phi(x)) G(x, y). \quad (1)$$

- (b) We would like to solve this differential equation. As a warmup, consider the case $g = 0$. Here is a trick: add a fictitious additional time direction T

$$(\partial_T - (\partial^2 + m^2)) G(x, y, T) = \mathbf{i}\delta^{d+1}(x-y)\delta(T) \quad (2)$$

This is just a diffusion equation (in $d+2$ dimensions and with a funny factor of \mathbf{i} !). Show that given a solution to (9), you can find the solution of (8) with $g = 0$ by

$$G(x, y) = \int_0^\infty dT G(x, y, T). \quad (3)$$

- (c) Show that the solution to the diffusion equation (9) is

$$G(x, y, T) = \frac{\mathbf{i}}{(2\pi T)^\alpha} e^{a\frac{(x-y)^2}{2T} + b\frac{m^2}{2}T}. \quad (4)$$

Find α, a, b . Use this to construct the path integral representation

$$G(x, y, T) = \int_{x(0)=x}^{x(T)=y} [Dx] e^{-\mathbf{i} \int_0^T d\tau (\dot{x}^\mu \dot{x}_\mu + m^2)}.$$

- (d) For the case of constant m^2 , the infinitesimal solution (12) actually works for finite T . Show by differentiation that plugging (12) into (10) gives an integral representation of the free Klein-Gordon propagator.
- (e) Now let $g \neq 0$ and suppose that ϕ is slowly varying. Generalize the path integral representation to include the dependence on ϕ .
- (f) Consider a non-relativistic situation, where the spacetime points x and y are separated by a timelike distance large compared to $1/m$. Justify and use stationary-phase methods to show that the dominant contribution to the path integral is a straight-line trajectory between the two points x and y . Evaluate the resulting amplitude as a functional of $\phi(x)$.

This calculation shows that the heavy particle made by the field Φ can be treated as a source for ϕ propagating on a fixed path in spacetime.

- (g) Redo the problem for a *charged* scalar field, Φ in the background of a vector potential A_μ , with

$$S[\Phi] = \int d^{d+1}x \frac{1}{2} (D_\mu \Phi^* D^\mu \Phi - m^2 \Phi^* \Phi), \quad D_\mu \Phi \equiv \partial_\mu \Phi - i A_\mu \Phi.$$

It will help to recall that the action of a classical charged particle is $\int d\tau (\dot{x}^2 + \dot{x}^\mu A_\mu(x))$.