

# Physics 215B QFT Winter 2025

## Assignment 3 – Solutions

Due 11:59pm Tuesday, January 28, 2025

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### 1. Brain-warmer.

Use the Clifford algebra to show that in 3+1 dimensions

$$\gamma^\mu (x\not{p} + m) \gamma_\mu = -2x\not{p} + 4m$$

where as usual  $\not{p} \equiv p^\mu \gamma_\mu$ . This identity will be useful in the numerator of the electron self-energy.

### 2. An example of renormalization in classical physics.

Consider a classical scalar field in  $D + 2$  spacetime dimensions coupled to an *impurity* (or defect or brane) in  $D$  dimensions, located at  $X = (x^\mu, 0, 0)$ . Suppose the field has a self-interaction which is localized on the defect. For definiteness and calculability, we'll consider the simple (quadratic) action

$$S[\phi] = \int d^{D+2}X \left( \frac{1}{2} \partial_M \phi(X) \partial^M \phi(X) + \frac{1}{2} g \delta^2(\vec{x}_\perp) \phi^2(X) \right).$$

Here  $X^M = (x^\mu, x_\perp^i)$ ,  $\mu = 0..D - 1$ ,  $i = 1, 2$ , *i.e.*  $x_\perp$  are coordinates transverse to the impurity.

Note, I changed the definition of  $g$  relative to the statement of the problem in order to avoid a proliferation of factors of 2.

This example is from [this paper](#) by Goldberger and Wise.

- What is the mass dimension of the coupling  $g$ ? This is why I picked a codimension<sup>1</sup>-two defect.
- Find the equation of motion for  $\phi$ . Where have you seen an equation like this before?

It's the Schrödinger equation for a particle in a 2d delta function potential.

- We will study the propagator for the field in a mixed representation:

$$G_k(x, y) \equiv \langle \phi(k, x) \phi(-k, y) \rangle = \int d^D z e^{i k_\mu z^\mu} \langle \phi(z, x) \phi(0, y) \rangle$$

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<sup>1</sup>An object whose position requires specification of  $p$  coordinates has codimension  $p$ .

– *i.e.* we go to momentum space in the directions in which translation symmetry is preserved by the defect. Find and evaluate the diagrams contributing to  $G_k(x, y)$  in terms of the free propagator  $D_k(x, y) \equiv \langle \phi(k, x) \phi(-k, y) \rangle_{g=0}$ . (We will not need the full form of  $D_k(x, y)$ .) Note that there are no loop diagrams, and in this sense, all the physics here is classical. Sum the series. I found it convenient to do this problem in Euclidean spacetime, so  $G$  and  $D$  are Euclidean propagators.

The euclidean path integral is of the form  $\int D\phi e^{-S_0} e^{-V}$  where  $S_0$  is the kinetic term and  $V = \int d^{D+2}x \delta^2(x_\perp) \frac{1}{2} g \phi^2$ . If we work in real time, the interaction vertex would be a factor of  $-i g \delta^{(2)}(x)$ . If we work in euclidean time, the two-point vertex is  $-g \delta^{(2)}(x)$ , and no  $i$ s will appear. From the sum of diagrams of the form (just as if we had done perturbation theory in the mass)

$$- \quad + \quad -x- \quad + \quad -x-x- \quad + \quad -x-x-x- \quad \dots$$

we find a geometric series

$$\begin{aligned} G_k(x, y) &= D_k(x, y) - g \int d^2 z_1 D_k(x, z_1) \delta^{(2)}(z_1) D_k(z_1, y) \\ &\quad + (-g)^2 \int d^2 z_1 \int d^2 z_2 D_k(x, z_1) \delta^{(2)}(z_1) D_k(z_1, z_2) \delta^{(2)}(z_2) D_k(z_2, y) + \dots \\ &= D_k(x, y) - g D_k(x, 0) D_k(0, y) + (-g)^2 D_k(x, 0) D_k(0, 0) D_k(0, y) \\ &\quad + (-g)^3 D_k(x, 0) D_k(0, 0)^2 D_k(0, y) + \dots \\ &= D_k(x, y) - g D_k(x, 0) (1 - g D_k(0, 0) + (-g)^2 D_k(0, 0)^2 + \dots) D_k(0, y) \\ &= D_k(x, y) - \frac{g}{1 + g D_k(0, 0)} D_k(x, 0) D_k(0, y). \end{aligned}$$

- (d) You should find that your answer to part 2c depends on  $D_k(0, 0)$ , which is divergent. This divergence arises from the fact that we are treating the defect as infinitely thin, as a pointlike object – the  $\delta^2$ -function in the interaction involves arbitrarily short wavelengths. In general, as usual, we must really be agnostic about the short-distance structure of things. To reflect this, we introduce a regulator. For example, we can replace the fourier representation of  $D_k(0, 0)$  with the cutoff version

$$D_k(0, 0; \Lambda) = \int_0^\Lambda d^2 q \frac{e^{i q \cdot 0}}{k^2 + q^2}. \quad (1)$$

Do the integral.

Note that the formula (1) would need an extra factor of  $\mathbf{i}$  if we were working in real time (in which case the interaction vertex would be  $-\mathbf{i}g\delta^2(x)$ , and the  $\mathbf{i}$ s would eat each other).

$$D_k(0, 0; \Lambda) = \int_0^\Lambda \mathrm{d}^2q \frac{e^{iq \cdot 0}}{k^2 + q^2} = \frac{1}{4\pi} \log \frac{\Lambda^2 + k^2}{k^2} \stackrel{\Lambda \gg k}{\cong} \frac{1}{4\pi} \log \frac{\Lambda^2}{k^2}.$$

These dimensions we're integrating here are spacelike, so there's no need for any Wick rotation.

- (e) Now we renormalize. We will let the *bare coupling*  $g$  (the one which appears in the Lagrangian, and in the series from part 2c) depend on the cutoff  $g = g(\Lambda)$ . We wish to eliminate  $g(\Lambda)$  in our expressions in favor of some measurable quantity. To do this, we impose a renormalization condition: choose some reference scale  $\mu$ , and demand that<sup>2</sup>

$$G_\mu(x, y) \stackrel{!}{=} D_\mu(x, y) - g(\mu)D_\mu(x, 0)D_\mu(0, y). \quad (2)$$

This equation defines  $g(\mu)$ , which we regard as a physical quantity. Show that (2) is satisfied if we let the bare coupling be  $g(\Lambda) = g(\mu)Z$ , with

$$Z = \frac{1}{1 - \frac{g(\mu)}{4\pi} \ln \left( \frac{\Lambda^2}{\mu^2} \right)}.$$

- (f) Find the beta function for  $g$ ,

$$\beta_g(g) \equiv \mu \frac{dg(\mu)}{d\mu},$$

and solve the resulting RG equation for  $g(\mu)$  in terms of some initial condition  $g(\mu_0)$ . Does the coupling get weaker or stronger in the UV?

You may be bothered that we previously defined the beta function as  $\Lambda \partial_\Lambda g(\Lambda)$ , in terms of the cutoff dependence. In a classically scale-invariant theory, the dependence on  $\Lambda$  and  $\mu$  is very closely tied together, since there are no other scales in the problem.

Solving for  $g(\Lambda)$  gives

$$g(\Lambda) = \frac{g(\mu)}{1 - \frac{g(\mu)}{4\pi} \log \frac{\Lambda^2}{\mu^2}}.$$

Then

$$\beta_g(g) = \frac{g(\Lambda)}{\left(1 - \frac{g(\Lambda)}{4\pi} \log \frac{\Lambda^2}{\mu^2}\right)^2} \frac{g(\Lambda)}{2\pi} = \frac{g^2(\mu)}{2\pi} = \frac{g^2}{2\pi}.$$

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<sup>2</sup>Note that if we worked in real time, there would be an extra  $\mathbf{i}$  in front of the second term on the RHS.

The solution is

$$g(\mu) = \frac{g(\mu_0)}{1 - \frac{g(\mu_0)}{2\pi} \log \frac{\mu}{\mu_0}}$$

which grows with  $\mu$ . Something bad happens when the denominator vanishes:

$$1 = \frac{g(\mu_0)}{2\pi} \log \frac{\mu_\star}{\mu_0}.$$

This scale  $\mu_\star$  where the coupling blows up is called a *Landau pole*.

3. **Scale invariance in QFT in  $D = 0 + 0$ , part 1.** [I got this problem from Frederik Denef.]

A nice realization of QFT in  $0 + 0$  dimensions is the statistical mechanics of a collection of non-interacting particles. The canonical partition function for a single particle (moving in one dimension) is

$$Z = \int dP dX e^{-\beta H} \propto \sqrt{T} Z_V(T) \quad (3)$$

with  $H = \frac{P^2}{2} + V(X)$  and  $T = 1/\beta$ . The momentum integral is Gaussian and we can just do it. The partition function of  $N$  non-interacting indistinguishable particles is then  $Z^N/N!$ , which just multiplies the energy  $U = T^2 \partial_T \log Z$  by a factor of  $N$ , so we don't miss anything by focussing on the single particle.

Let's consider the case

$$V(X) = aX^2 + bX^4 + cX^6 \quad (4)$$

and figure out the important features of the temperature dependence of the thermodynamic quantities by scaling arguments.

- (a) Assuming  $a \neq 0, b \neq 0, c \neq 0$ , find the behavior of the thermal energy  $U$  and the heat capacity  $C = \partial_T U$  in the limit  $T \rightarrow 0$  and in the limit  $T \rightarrow \infty$  using scaling arguments. Which parts of the potential determine the respective limiting behaviors?

First, to understand the low-temperature behavior, let  $x \equiv X/\sqrt{T}$ , so that

$$Z_V = \int dX e^{-V(X)/T} = T^{1/2} \int dx e^{-(ax^2 + bx^4T + cx^6T^2)} = T^{1/2} \underbrace{\int dx e^{-ax^2}}_{\text{some number}} \underbrace{e^{-(bx^4T + cx^6T^2)}}_{\xrightarrow{T \rightarrow 0} 1}. \quad (5)$$

Therefore,  $Z \stackrel{T \rightarrow 0}{\propto} T^{1/2+1/2}$ . In this case the quadratic term is most important. If  $Z \propto T^\alpha$  then  $U = \alpha T$  and  $C = \alpha$ , so here  $C = 1$ . To understand the high-temperature behavior, let  $y \equiv X/T^{1/6}$  so that

$$Z_V = \int dX e^{-V(X)/T} = T^{1/6} \int dy e^{-(cy^6 + ax^2/T^{2/3} + bx^4/T^{2/3})} = T^{1/6} \int dy e^{-cy^6} \underbrace{e^{-(ax^2/T^{2/3} + bx^4/T^{2/3})}}_{T \rightarrow \infty \rightarrow 1}. \quad (6)$$

So at high temperatures  $C \rightarrow \frac{1}{2} + \frac{1}{6}$ . At high temperature, the particle can explore the whole potential and the highest power in the potential is what matters.

- (b) If some of the couplings  $a, b, c$  vanish, the low or high temperature scaling behavior may change. For example, what is the heat capacity at low temperature when  $a = 0, b \neq 0$ ?

In this case, the quartic term dominates and  $Z_V \sim T^{1/4}$  and  $C = 3/4$ .

A word about notation: the symbol  $\sim$  is often used by physicists to indicate a scaling relationship, where the constant prefactors are neglected. The relation we derive here for  $C$  however is an equality in the relevant regime of temperatures – the constant is the thing that matters.

- (c) When  $b$  is sufficiently large (and  $a \neq 0, c \neq 0$ ), there will be an intermediate temperature regime over which the heat capacity is again constant, but different from the low- and high-temperature limits. What is this heat capacity?

Same as the previous part.

- (d) In general, we can think of the change of  $C$  with  $T$  as a kind of classical renormalization group (RG) flow, interpolating between ‘fixed points’ where  $C$  becomes constant. In general, these fixed points correspond to potentials  $V(X)$  with a scaling symmetry  $V(\lambda^\Delta X) = \lambda V(X)$  for some choice of scaling dimension  $\Delta$  of  $X$ . What is the heat capacity for a fixed point with scaling dimension  $\Delta$  for  $X$ ?

$$Z_V = \int dX e^{-V(X)/T} = \int dX e^{-V(T^{-\Delta} X)} = T^\Delta \underbrace{\int dx e^{-V(x)}}_{\text{indep of } T} \quad (7)$$

with  $x \equiv T^{-\Delta} X$ . So  $Z \propto T^{1/2+\Delta}$  and

$$C = \Delta + \frac{1}{2}. \quad (8)$$

- (e) Borrowing more language of the renormalization group, we can classify deformations  $\delta V(X) = \epsilon X^m$  of a fixed point  $V(X) \propto X^{2n}$  as irrelevant, marginal, or relevant, depending on whether the deformation becomes dominant or negligible in the IR limit, *i.e.* in the limit of low  $T$ . Here and below  $\epsilon$  can take on any value, not necessarily small. Restricting to deformations with an  $X \rightarrow -X$  symmetry, what are the relevant and irrelevant deformations of  $V(X) = X^{2n}$ ? (Note that a deformation  $\delta V = \epsilon X^{2n}$  can be absorbed into a redefinition of  $X$ , which does not change the heat capacity.) Lower powers than  $2n$  are relevant, higher powers are irrelevant.

- (f) The  $T$ -dependence of correlation functions (here, expectation values of powers of  $X$ ) at fixed points is also determined by the scaling properties. What is the  $T$ -dependence of  $\langle X^k \rangle$  at a fixed point  $V(X) = X^{2n}$ ?

$$\langle X^k \rangle = \frac{\int dX X^k e^{-V(X)/T}}{Z_V} = \frac{T^{\Delta(1+k)} \int dx x^k e^{-V(x)}}{T^\Delta} \propto T^{\Delta k}$$

where  $\Delta = \frac{1}{2n}$  is the scaling dimension of  $X$ .

- (g) Non-polynomial  $V(X)$  can be considered as well. For example, what is the heat capacity at small and large  $T$  for  $V(X) = (1 + X^2)^{1/n}$ ?

Since this function still grows at large  $X$ , the high-temperature behavior is dominated by the large- $X$  behavior where  $V(X) \sim X^{2/n}$ , so  $\Delta = n/2$  and  $C = \frac{n+1}{2}$ . At low temperature, we Taylor expand in small  $X$  to find  $V(X) \sim 1 + X^2/n$  and find  $\Delta = 1/2$  and  $C = 1$ , where we used (8).